
Subject: Re: Latitude longitude and Image Navigation
Posted by [AISHWARYA](#) on Fri, 24 Jun 2011 18:51:08 GMT
[View Forum Message](#) <> [Reply to Message](#)

On Jun 24, 10:08 pm, Brian Wolven <brian.wol...@gmail.com> wrote:

> As I recall, I was generating a lat-long grid to overlay on HST/STIS images of Galilean satellites. It looks like I used the North Pole position vector and the sub-observer point (calculated in SPICE) to get the viewing orientation and then drew the projected lat-long grid based on that information. This was a long time ago (15 years?), albeit in galaxy not so far away. I think the code below is what I used to calculate the rotation matrix for the grid; I don't know if it will help or not. Looks like I first rotated the z-axis to be aligned with the pole vector, then found the rotation about the pole to place the sub-observer point at the proper coordinates.

```
>
> function polar_rotation,body
> ;=====
=====
> ; Rotate z-axis into specified pole position vector (nppv)
> ;=====
=====
> nppv = fourvector(unit_vector(body.nppv)) ; unit 4-vector along NPPV
> zaxis = [0.,0.,1.,0.]
> rpol1 = rotate_a2b(zaxis,nppv)
> ;=====
=====
> ; Latitude must be converted from planetographic to planetocentric coordinates.
> ; Planetocentric co-latitude is equivalent to theta in spherical coordinates.
> ; -1 * longitude is equivalent to phi in spherical coordinates.
> ;=====
=====
> axisratio = body.r_polar/body.r_equatorial
> sublatg = body.subobsrlat
> sublatc = pg2pc_lat(sublatg,axisratio)
> subcolat = (90.-sublatc)*!dtr ; theta in spherical coords
> sublon = -1.*body.subobsrlon*!dtr ; phi in spherical coords
> ;=====
=====
> ; Rotate body such that line of sight intersects specified subobserver point.
> ; Rotate only in the plane perpendicular to the pole (nppv)
> ;=====
=====
> subdes = [cos(sublon)*sin(subcolat),sin(sublon)*sin(subcolat),cos(sub colat)]
> subcur = -1.*body.pos/length(body.pos) ;
> oldang = angle_between(subdes,zaxis)
> bodang = angle_between(subcur,nppv)
> subdes = fourvector(subdes)#rpol1 ;4-vector of desired sub-observer pt
> subcur = fourvector(subcur) ;4-vector of initial sub-observer pt
> newang = angle_between(subdes,nppv)
> newdes = subdes - nppv*dot_product(nppv,subdes) ; Project components on plane
```

```

> newsub = subcur - nppv*dot_product(nppv,subcur) ; perpendicular to nppv.
> rpol2 = rotate_a2b(newdes,newsub)
> finang = angle_between(subdes#rpol2,nppv)
> ;=====
=====
> ; Diagnostics
> ;=====
=====
> ; print,'Curr: ',float(reform(subcur))
> ; print,'Dsrd: ',float(reform(subdes))
> ; print,float(reform([sublon!/dtor,sublatg!/dtor,sublatc,subco lat!/dtor]))
> ; print,float([bodang,oldang,newang,finang])
> ;=====
=====
> ; Return total matrix to calling procedure.
> ;=====
=====
> return,rpol1#rpol2
> end

```

Thank you, i'm trying to understand the code. But, is there any way that I can measure the pixel corresponding to North pole angle and distance by any inbuilt function? Does plate scale come into picture anywhere?

Method 2 : I was trying my hands on Map_Image. I could fix the centre but lat min and Lat max seems to be a question. Do you think Map_image can help me in this?

Also, I recently found a SOHO code written for converting camera coordinates to solar coordinates by Dr. Iwang

this is the link : http://www.astro.washington.edu/docs/idl/cgi-bin/getpro/library32.html?FIND_LIMB2

I was trying a similar one but IDL stops running when the first while loop is encountered.

Regards,
Aishwarya.
