
Subject: Re: flux-conserving image resampling?

Posted by [Bringfried Stecklum](#) on Thu, 23 Feb 2006 08:32:09 GMT

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Marshall Perrin wrote:

> Some time ago, there was a thread in this newsgroup on 'drizzle'-like
> image resampling methods, much of which is preserved on David's web
> page. Well, actually it was mostly a thread on clever and
> convoluted histogram tricks, and resulted in some speedy solutions to
> array decimation. But the ultimate consensus then was that there is
> no good drizzle-like flux conserving image resampler in IDL, and maybe
> C is a better way to go.

>
> Before I sit down and reinvent the wheel on this one, I thought I'd
> ask around and see what, if any, code for this might exist now a few
> years later. I'm looking for a good way to take pixellated data
> measured in a camera with known field distortion and resample them
> onto a regular grid. Drizzle or PyDrizzle would do what I want,
> except it seems fairly nontrivial to get them to accept non-HST data.
> (Am I wrong? If anyone has tips on how to do this, I'm all ears.)
> If anyone knows of an IDL solution for this, or a C DLM, I'd be most
> appreciative... And failing that, at least knowing definitively that
> no such thing exists now would probably help fortify my spirits for
> writing my own!

>
> Thanks,
>
> - Marshall
>
>

Dear Marshall,

I am not quite sure whether you really need a drizzle algorithm for correcting the image distortion. In IRAF, the `geotran` task does geometrical image transformation. The necessary distortion coefficients can be derived with `geomap`. Here are a few lines of IDL code I've written to correct the image distortion of ESO's ISAAC NIR instrument using 2nd order polynomials. I did not check how well it preserves the flux. In any case, you can force the total flux of the output image to be the same as that of the input image.

Regards,

Bringfried

;-----

```

function cdistort,image,px=px,py=py,resample=resample

; distortion coefficients, assume ISAAC as default
if not keyword_set(px) then px=[0.7246,-0.3609,-0.04242,-0.05569,1.905,1.887,-0.0564]
if not keyword_set(py) then py=[-0.3529,0.1274,0.8327,2.411,0.004007,0.1202,2.386]

if not keyword_set(resample) then resample=1

npix=resample*1024

x=findgen(npix)
xni=(2*(x-npix/2.)/npix)#replicate(1.,npix)
yni=transpose(xni)
xni2=xni^2
yni2=yni^2

dx=resample*(px[0]*xni*yni+px[1]*xni2+px[2]*yni2+px[3]*xni2*
yni+px[4]*xni*yni2+px[5]*xni2*xni+px[6]*yni2*yni)
dy=resample*(py[0]*xni*yni+py[1]*xni2+py[2]*yni2+py[3]*xni2*
yni+py[4]*xni*yni2+py[5]*xni2*xni+py[6]*yni2*yni)

if not keyword_set(resample) then cimage=image else
cimage=CONGRID(image,npix,npix,cubic=-0.5)

xni=(xni+1)*npix/2.
yni=(yni+1)*npix/2.
xci=xni+dx
yci=yni+dy

if not keyword_set(resample) then return,cimage[xci,yci] else
return,rebin(cimage[xci,yci],1024,1024)

end

```

File Attachments

1) [stecklum.vcf](#), downloaded 103 times

Subject: Re: flux-conserving image resampling?
 Posted by [henrygroe](#) on Thu, 23 Feb 2006 16:00:02 GMT
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I've faced similar problems; The thing I would worry about with some of the 'flux conserving' algorithms out there is that while the total flux in the image is conserved, that does not mean the total flux in each object is conserved.

-Henry

Subject: Re: flux-conserving image resampling?

Posted by [Marshall Perrin](#) on Thu, 23 Feb 2006 17:39:19 GMT

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Bringfried Stecklum <stecklum@tls-tautenburg.de> wrote:

> I did not check how well it preserves the flux. In any case, you can
> force the total flux of the output image to be the same as that of the
> input image.

As Henry points out, this isn't really sufficient to ensure that the original flux stays in the right parts of the final image. That's important for my particular application, which is a dual-beam differential polarimeter. I'm trying to take a very precise difference between two images of the same star taken simultaneously on different parts of the detector, and thus merely enforcing total flux conservation lets light slop between the two stellar images (and the background too). Yes, it's a small effect, but I'm trying to -measure- a fairly small difference in the first place and thus care about this.

- Marshall

Subject: Re: flux-conserving image resampling?

Posted by [Bringfried Stecklum](#) on Fri, 24 Feb 2006 09:03:13 GMT

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Marshall Perrin wrote:

> Bringfried Stecklum <stecklum@tls-tautenburg.de> wrote:

>
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> but I'm trying to -measure- a fairly small difference in the first place and
> thus care about this.

>
> - Marshall

Ah, I see your point. I guess what has to be done is to account for the difference in pixel size between the original and the corrected image.

The code I've written only does resampling on a grid which represents the pixel location of the undistorted image. In order to preserve the flux locally, one has to keep in mind that depending on the distortion one pixel of the corrected image will subtend an area of less or more than one pixel in the original image. Well, I'm not talking about sub-pixels here (we had this earlier in the group...). For instance in the simple case of a Wollaston which has slightly different image scales for the O and E beams, the ratio of those scales should be applied after one image has been registered to the other. In some sense this photometric correction can be considered as distortion flatfield.

- Bringfried

Subject: Re: flux-conserving image resampling?

Posted by edward.s.meinel@aero on Fri, 24 Feb 2006 16:04:14 GMT

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If you are interested in the flux difference of two images of the same star, can't you just calculate the total flux of each star and then subtract the results? I assume you aren't looking at the sun, so the star would be unresolved...

Ed

Marshall Perrin wrote:

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